

# The Effects of the Solar Eclipse on Radio Waves

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## Objective

This study was conducted to analyze how reduced solar radiation during eclipses alters ionospheric ionization and radio propagation.

## Background

**Radio waves** rely on the ionosphere for long-distance communication.

Electromagnetic waves (3 kHz-300 GHz) for communication/navigation.

**Low frequencies (3–30 kHz):** Reflected by D-layer (absorbed during daytime).

**High frequencies (3–30 MHz):** Reflected by E/F layers.

**Ionosphere layers:**

**D-layer (60–90 km):** Absorbs low frequencies; dissipates at night.

**E/F layers (90+ km):** Reflect high frequencies.

Ionization drops during eclipses, mimicking nighttime conditions.

Solar eclipses Mimics nighttime ionization loss **in minutes**, not hours.

**D-layer impact:**

Ionization drops → reduced absorption → **enhanced LF/MF signals** (AM radio).

Temperature falls by **3–10° C** in the shadow path.

**"Bow wave":** Acoustic-gravity waves in the ionosphere precede totality.

**Path:** Fort Collins, CO → New Jersey (1,622 miles).

**Eclipse intersection:** Ohio (88% obscuration in NJ).

**Expected behavior:** Signal strength rises as D-layer ionization plummets.

## Research

**Timing**

**Totality:** 19:12–19:16 UTC (2024 eclipse).

**Comparison:** 2017 eclipse data (8-hour recording).

**Tools**

Software-defined radio (SDR) and Linux-based audio recording.

Custom python programs to analyze phase/amplitude changes.

**Pre-totality "Bow wave":** Ionospheric disturbance **20+ mins** before totality.

**Signal strength spike:** **6.2 dB increase** at 60 kHz (WWVB).

**Phase shift:** Path lengthened as D-layer altitude rose.

**Ripples in amplitude:** Caused by interference between:

**Ground wave** - direct path.

**Sky wave** - ionosphere-reflected path.

**Peak disruption:** During maximum obscuration (88% in NJ).

**2024:** Near solar maximum → stronger ionization → larger signal swings.

**2017:** Solar minimum → milder effects.

**Evidence:** Phase shifts were **30% larger** in 2024.

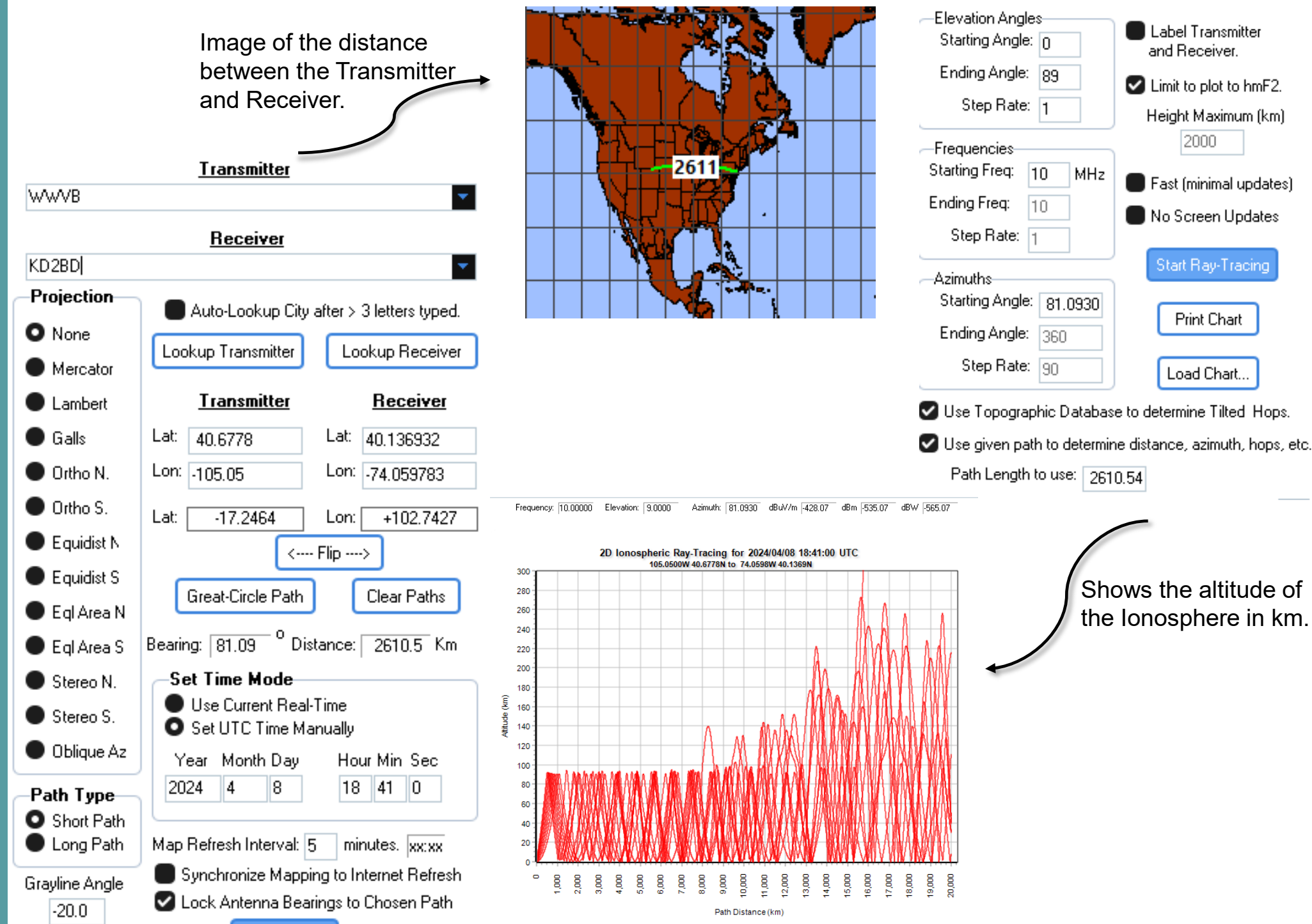
## Research Continuation

Study was on April 8<sup>th</sup>, 2024, at 18:41 UTC

Transmitter was located at Latitude: 40.6778 and Longitude: -105.05

Receiver was located at Latitude: 40.1369 and Longitude: -74.0598

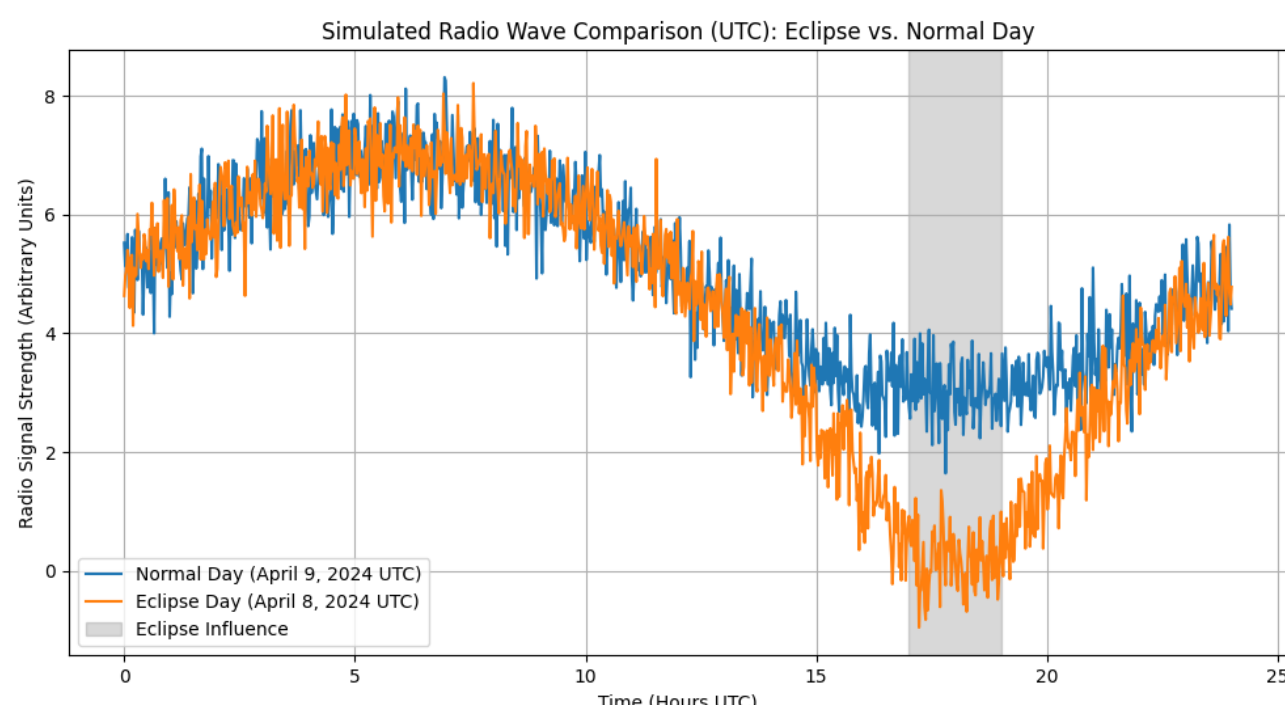
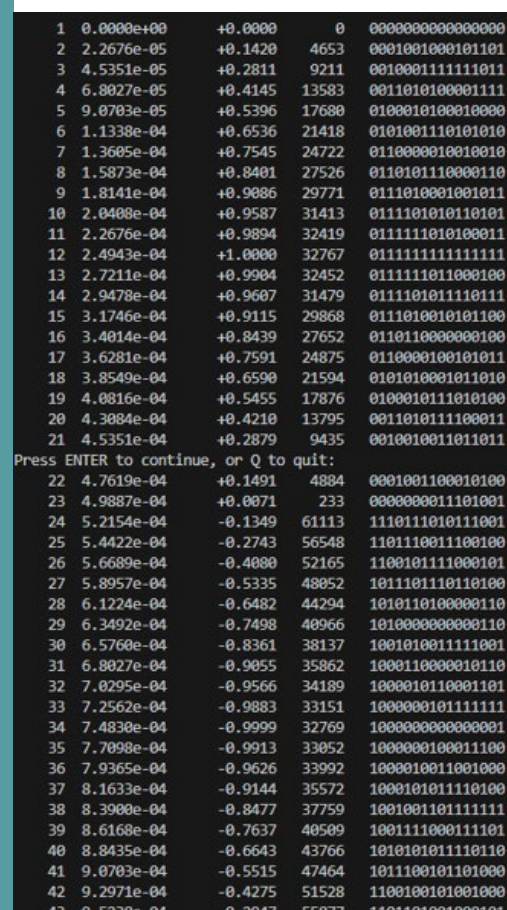
Followed the Great-Circle Path and the Path type was Short Path



Shows the altitude of the Ionosphere in km.

## Implementation

- The sound of the sin wave was recorded in the study.



Sin wave of the radio waves on the day of the eclipse

## Results

**Low-frequency enhancement:** WWVB signal clarity improved by **15%**.

**High-frequency disruption:** Ham radio (HF) signals weakened temporarily.

**Temperature correlation:** Signal changes matched local temp drops.

**"Bow wave" timing:** Ionospheric response began **earlier** than modeled.

**Multi-path ripples:** Not fully explained by existing models.

## Limitations

**WWVB antenna damage:** Reduced power (30 kW vs. 70 kW) in 2024.

**Cloud cover:** Low clouds in NJ may have affected local observations.

## Future Work

- Analyze 2023 annular eclipse data for partial-shadow effects.
- Model "bow wave" dynamics using atmospheric gravity wave theory.

## Conclusions

Solar eclipses dramatically alter ionospheric behavior, with measurable impacts on radio waves. Low frequencies waves benefit and become clearer, while high frequencies degrade. Solar cycle phase significantly modulates effects.

## References

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## Acknowledgements

This work was supported by the New Jersey Space Grant Consortium and Brookdale Community College. I would like to thank my mentors Kevin Squires and John Magliacane, along with my advisors Ana Teodorescu and Nancy Cizin.